## Excited states above the Hoyle state

## Y. Funaki<sup>\*1</sup>

In <sup>12</sup>C, there exists besides the Hoyle state (the second 0<sup>+</sup> state of <sup>12</sup>C) a number of other  $\alpha$  gas states above the Hoyle state that one can qualify as excited states of the Hoyle state. For a description of those states, it is useful to adopt a generalized THSR wave function<sup>1)</sup>, as the so-called THSR wave function<sup>2,3)</sup> is known to be the best to describe the Hoyle state<sup>4)</sup>. The part of the  $3\alpha$  THSR wave function that contains the c.o.m. motion of the  $\alpha$  particles contains two Jacobi coordinates  $\boldsymbol{\xi}_1$  and  $\boldsymbol{\xi}_2$ . As a natural extension of the original THSR wave function, it is possible to associate two different width parameters  $B_1, B_2$  with the two Jacobi coordinates. In this case the translationally invariant THSR wave function has the following form:

$$\Psi_{3\alpha}^{\text{thsr}} = \mathcal{A} \bigg[ \exp \bigg( -\frac{4}{3B_1^2} \boldsymbol{\xi}_1^2 - \frac{1}{B_2^2} \boldsymbol{\xi}_2^2 \bigg) \phi_1 \phi_2 \phi_3 \bigg].$$
(1)

With this type of generalized THSR wave function, one can get a much richer spectrum of <sup>12</sup>C. Axial symmetry has been assumed and the four *B* parameters are taken as Hill-Wheeler coordinates. In Fig. 1, the calculated energy spectrum is shown. One can see that besides the ground state band, there are many  $J^{\pi}$  states obtained above the Hoyle state. All these states turn out to have large rms radii (3.7 ~ 4.7 fm), and therefore can be considered as excitations of the Hoyle state. The Hoyle state can thus be considered as the "ground state" of a new class of excited states in <sup>12</sup>C. In particular, the nature of the series of states  $(0^+_2, 2^+_2, 4^+_2)$ and the  $0^+_3$  and  $0^+_4$  states have recently been widely discussed from the experimental side<sup>5-9</sup>.

In Fig. 1, the E2 transition strengths between Jand  $J \pm 2$  states and monopole transitions between  $0^+$ states are shown with corresponding arrows. We note the very strong E2 transitions inside the Hoyle band,  $B(E2; 4_2^+ \to 2_2^+) = 591 \text{ e}^2 \text{fm}^4 \text{ and } B(E2; 2_2^+ \to 0_2^+) = 295 \text{ e}^2 \text{fm}^4$ . The transition between the  $2_2^+$  and  $0_3^+$ states is also very large,  $B(E2; 2_2^+ \rightarrow 0_3^+) = 104 \text{ e}^2 \text{fm}^4$ . There have been attempts to interpret the Hoyle band as a rotational band of a spinning triangle as this was successfully done for the ground state band<sup>10</sup>). However, the situation may not be as straightforward as it seems<sup>11)</sup>. Since the two transitions  $2^+_2 \rightarrow 0^+_2$  and  $2^+_2 \rightarrow 0^+_3$  are of similar magnitude, no clear band head can be identified. Whether they can be qualified as members of a rotational band or of a vibrational band or a mixture of both is an open question. One should also realize that the  $0^+_3$  state is strongly excited from the Hoyle state by monopole transition whose strength is obtained from the extended THSR calculation to be  $M(E0; 0_3^+ \to 0_2^+) = 35 \text{ fm}^2$ . Therefore, the  $0_3^+$  state



Fig. 1. Spectrum of <sup>12</sup>C obtained from the extended THSR approach in comparison with experiment.

seems to be a state where one  $\alpha$  particle has been lifted out of the condensate to the next higher S level with a node. This is confirmed by the  $S^2$  factor, which is calculted to be close to unity for the  ${}^{12}C(0^+_1) + \alpha(S)$  channel. We can also conclude that the  $2^+_3$  and  $4^+_3$  states, together with the  $0^+_3$  state, form the higher nodal rotational band, from the Reduced Width Amplitude analysis. On the other hand, the  $0^+_4$ ,  $2^+_4$  and  $4^+_4$  states all have the largest contribution from the  ${}^{8}Be(2^+) + \alpha(D)$ channel, indicating that these states are built out of an  $\alpha$  particle orbiting in a D-wave around a (correlated) two  $\alpha$  pair, also in a relative 0D state.

References

- 1) Y. Funaki et al., Prog. Part. Nucl. Phys. 82, 78 (2015).
- 2) A. Tohsaki et al., Phys. Rev. Lett. 87, 192501 (2001).
- 3) Y. Funaki et al., Prog. Theor. Phys. 108, 297 (2002).
- 4) Y. Funaki et al., Phys. Rev. C 67, 051306(R) (2003).
- 5) M. Itoh *et al.*, Nucl. Phys. A **738**, 268 (2004).
- 6) M. Itoh et al., Phys. Rev. C 84, 054308 (2011).
- 7) M. Freer et al., Phys. Rev. C 80, 041303(R) (2009).
- W. R. Zimmerman *et al.*, Phys. Rev. Lett. **110**, 152502 (2013).
- 9) M. Freer et al., Phys. Rev. C 83, 034314 (2011).
- 10) D. J. Maŕin-Lámbarri *et al.*, Phys. Rev. Lett. **113**, 012502 (2014).
- 11) Y. Funaki, Phys. Rev. C 92, 021302 (2015).

<sup>\*1</sup> RIKEN Nishina Center