Shell-model study of nuclear structure around ¹⁰⁰Sn

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The nuclear structure around the doubly-magic N=Z nucleus ¹⁰⁰Sn has been of great interest from various viewpoints such as the development of shellstructure and the proton-neutron correlations. For a reliable prediction of unknown targets by the shell model, one of our strategies is to minimally modify so-called G-matrix interactions¹) by fitting the shellmodel results to available experimental energy data. In the previous work²⁾, we have determined an effective interaction called JUN45 in the model space covering nuclei with 28 < N, Z < 50. Also, we have tried the shell-model fits to describe Sn isotopes with $N{=}50 \sim$ 82 and obtained an effective interaction $SNBG1^{(3)}$. Since the ¹⁰⁰Sn is located at the end of the model space in both studies, it was impossible to discuss the excitation across the N and/or Z=50 shell closure. In this report, we present another approach along this line, aiming at the description of nuclei including $^{100}\mathrm{Sn}.$

We take four single-particle orbits $1p_{1/2}$, $0g_{9/2}$, $1d_{5/2}$ and $0g_{7/2}$ for both protons and neutrons assuming a hypothetical "core" $^{76}_{38}$ Sr₃₈. This choice is motivated by the excellent success of the $(p_{1/2}, q_{9/2})$ model space near the $N \sim 50$ lines due to the approximate degeneracy of these orbits around there, as suggested in Fig.1(a). Also, since the $7/2^+$ state comes down rapidly as the proton number is increased towards Z = 50 (see Fig.1(b)), the last two orbits $(d_{5/2}, g_{7/2})$ are essential. Based on the information about the dominant configurations obtained with the JUN45 and the SNBG1 interactions, we have selected the experimental data in the range of $47 \le N \le 58$ for the fit. In order to reduce the amount of computation for the fitting, we take the t=4 truncated model space, where t stands for the maximum number of nucleons that can excite from the $(p_{1/2}, g_{9/2})$ orbits to the $(d_{5/2}, g_{7/2})$ orbits relative to the naive lowest configuration. Starting from the G-matrix interaction derived from the N³LO interaction⁴), we have carried out a series of iterative fits. We assume the isospin symmetry, and adopt the $A^{-0.3}$ mass-dependence of the two-body matrix element (TBME). In the latest fit, 197 TBMEs and 4 single-particle energies have been determined with a rms error of 231keV for 528 data.

As examples of the fitted results, the energy levels of

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low-lying states are shown in Fig.1 for odd-mass isotones with N = 50 and 51. It can be seen that the overall trends are reasonably described by the present shell-model calculations. As for ¹⁰⁰Sn, using this interaction at the t = 6 truncation level, the excitation energy of the 2_1^+ state is predicted to be 4.8MeV, and the 0p-0h component in the ground-state wavefunction is 71%. The calculated $B(E2; 0^+ \rightarrow 2^+)=0.13 e^2 b^2$ with the effective charges $e_p=1.5$, $e_n=0.5$ is almost consistent with the shell-model result in a different model space⁷⁾.



Fig. 1. Energy levels of low-lying states for (a) N=50 isotones with odd-number of protons and (b) N=51 isotones with even-number of protons. Calculated 1/2⁻, 9/2⁺, 5/2⁺ and 7/2⁺ states are shown with dashed, long-dashed, solid and dotted lines, respectively, which are compared with the experimental data denoted by diamonds, triangles, circles and squares, respectively. Experimental data are taken from Ref.⁵⁾, where uncertain spin assignments are explicitly shown. The shell-model results are obtained by using the efficient code MSHELL64⁶).

References

- 1) M. Hjorth-Jensen et al.: Phys. Rep. 261, 125 (1995).
- 2) M. Honma *et al.*: Phys. Rev. C **80**, 064323 (2009).
- 3) M. Honma et al.: RIKEN Accel. Prog. Rep. 45, 35 (2012).
- 4) D. R. Entem et al.: Phys. Rev. C 68, 041001(R) (2003).
- 5) Data extracted using the NNDC WorldWideWeb site from the ENSDF database.
- 6) T. Mizusaki et al.: MSHELL64 code (unpublished).
- 7) G. Guastalla et al.: Phys. Rev. Lett. 110, 172501 (2013).