

Evidence for a new nuclear ‘magic number’ in $^{54}\text{Ca}^\dagger$

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Over recent years, the evolution of nuclear shell structure in exotic, neutron-rich nuclei has attracted much attention on both the experimental and theoretical fronts. In the neutron-rich fp shell, the onset of the $N = 32$ subshell closure is well established from the structural characteristics of $^{52}\text{Ca}^{1,2)}$, $^{54}\text{Ti}^{3,4)}$ and $^{56}\text{Cr}^{5,6)}$. This subshell gap is reproduced successfully by numerous theoretical predictions. In the framework of tensor-force-driven shell evolution⁷⁾, the onset of the $N = 32$ subshell closure results as a direct consequence of a sizable $\nu p_{3/2}-\nu p_{1/2}$ gap, which presents itself as the $\nu f_{5/2}$ orbital shifts up in energy owing to a weakening of the attractive $\pi f_{7/2}-\nu f_{5/2}$ interaction as protons are removed from the $\pi f_{7/2}$ orbital. Another important manifestation of some theories is the prediction of a large subshell gap at $N = 34$, which develops if the $\nu f_{5/2}$ orbital lies sufficiently high in energy above the $\nu p_{1/2}$ orbital. It has already been shown that no significant $N = 34$ subshell gap exists in $^{56}\text{Ti}^{4,8)}$ or $^{58}\text{Cr}^{6,9)}$ and, therefore, the size of the energy gap in ^{54}Ca is an important structural characteristic that requires experimental input. Moreover, the single-particle states of ^{53}Ca should also reflect the nature of the $N = 34$ subshell closure in isotopes far from stability.

The structures of ^{54}Ca and ^{53}Ca were investigated using in-beam γ -ray spectroscopy at the RIBF to address this issue. A primary beam of $^{70}\text{Zn}^{30+}$ ions at 345 MeV/u was used to create a radioactive beam containing ^{55}Sc and ^{56}Ti , which was focused on a 10-mm-thick

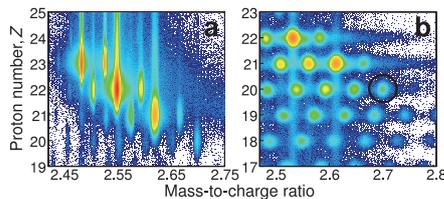


Fig. 1. (colour) Particle identification plots measured by (a) the BigRIPS separator and (b) the ZeroDegree spectrometer. The black circle indicates ^{54}Ca events.

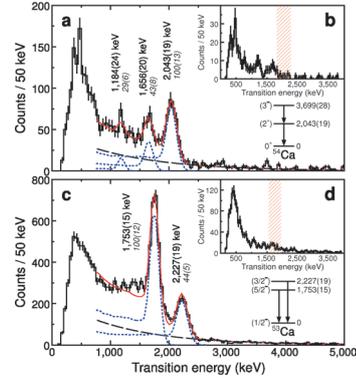


Fig. 2. (colour) Doppler-corrected γ -ray energy spectra for (a) ^{54}Ca and (c) ^{53}Ca . Insets (b) and (d) indicate γ rays in coincidence with the 2043- and 1753-keV lines.

Be reaction target located inside the DALI2 γ -ray detector array at F8. Reaction products were identified with the ZeroDegree spectrometer (see Fig. 1).

The energy spectra for ^{54}Ca and ^{53}Ca deduced in the present work are presented in Fig. 2. The most intense peak in the ^{54}Ca spectrum, the line at 2043(19) keV, is assigned as the $2_1^+ \rightarrow 0^+$ ground-state transition. Several other weaker lines are also reported. The relatively high energy of the 2_1^+ state reflects the doubly magic nature of ^{54}Ca and provides direct experimental evidence for the onset of a sizable subshell closure in $N = 34$ isotones far from stability. Shell-model calculations adopting a modified GXPF1B Hamiltonian indicate that the strength of the $N = 34$ subshell gap in ^{54}Ca (the $\nu p_{1/2}-\nu f_{5/2}$ SPO energy gap) is in fact comparable to the $N = 32$ subshell gap in ^{52}Ca (the $\nu p_{3/2}-\nu p_{1/2}$ SPO energy gap) (see original Letter for details). In the ^{53}Ca spectrum, the 1753(15)-keV transition is reported for the first time, while the line at 2227(19) keV is consistent in energy with a transition previously measured in a decay study¹⁰⁾.

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