

Three quasiparticle isomers in odd-even $^{159,161}\text{Pm}$: Calling for modified spin-orbit interaction for the neutron-rich region[†]

R. Yokoyama,^{*1,*2} E. Ideguchi,^{*3,*8} G. Simpson,^{*4} Mn. Tanaka,^{*3} Y. Sun,^{*5} C. -J. Lv,^{*5} Y. -X. Liu,^{*6} L. -J. Wang,^{*7} S. Nishimura,^{*8} P. Doornenbal,^{*8} G. Lorusso,^{*8} P. -A. Söderström,^{*8} Z. Y. Xu,^{*11,*2} J. Wu,^{*10,*8} T. Sumikama,^{*8} N. Aoi,^{*3,*8} H. Baba,^{*8} F. L. Bello-Garrote,^{*12} G. Benzoni,^{*13} F. Browne,^{*14,*8} R. Daido,^{*15} Y. Fang,^{*15} N. Fukuda,^{*8} A. Gottardo,^{*16,*17} G. Gey,^{*18,*8} S. Go,^{*1} S. Inabe,^{*8} T. Isobe,^{*8} D. Kameda,^{*8} K. Kobayashi,^{*19} M. Kobayashi,^{*1} I. Kojouharov,^{*20} T. Komatsubara,^{*21,*22} T. Kubo,^{*8} N. Kurz,^{*20} I. Kuzi,^{*23} Z. Li,^{*10} M. Matsushita,^{*1} S. Michimasa,^{*1} C. B. Moon,^{*24} H. Nishibata,^{*15} I. Nishizuka,^{*9} A. Odahara,^{*15} Z. Patel,^{*8,*25} S. Rice,^{*8,*25} E. Sahin,^{*12} H. Sakurai,^{*8,*11} H. Schaffner,^{*26,*8} L. Sinclair,^{*26,*8} H. Suzuki,^{*8} H. Takeda,^{*8} J. Taprogge,^{*27,*28} Zs. Vajta,^{*23} H. Watanabe,^{*29,*8} and A. Yagi^{*15}

Nuclear properties of neutron-rich rare-earth nuclei at $Z \approx 60$ could be the possible key to answer one of the longstanding astrophysical questions: the formation of the $A \approx 160$ (rare-earth) peak observed in the elemental abundance distribution. From the nuclear-structure perspective, the formation of a peak in elemental abundance originates in extra stability of local nuclei. For example, in r -process, the neutron magic number $N = 82$ and $N = 126$ are responsible for the prominent abundance peaks at $A \approx 130$ and $A \approx 195$, respectively. Mumpower *et al.* identified that the nuclear properties of nuclei at $N \approx 100$ are critical to the rare-earth peak formation.¹⁾ In the case of the deformed rare-earth nuclei, a large shell gap between Nilsson single-particle orbitals stabilizes the nuclear shape

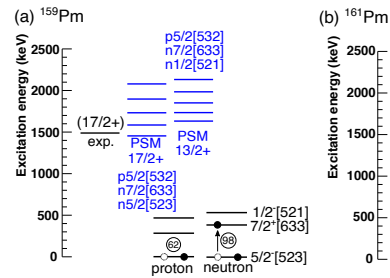


Fig. 1. Excitation energies of the isomers in (a) ^{159}Pm and (b) ^{161}Pm compared with the PSM calculation with the modified Nilsson parameters.

at large deformation. Several theoretical works have predicted the existence of the deformed shell gaps at $N \approx 100$.²⁾ Recently, a $6^- \nu(5/2[523] \otimes 7/2[633])$ state was discovered in $^{162}\text{Gd}_{98}$,³⁾ indicating the large shell gap at $N = 98$.

The spectroscopic study on neutron-rich rare-earth nuclei was performed at RIBF using the in-flight fission of $^{238}\text{U}^{86+}$ beam. The delayed γ rays were measured using a Ge detector array, EURICA.⁴⁾ Accordingly, we discovered new quasi-particle isomers in $^{159,161}\text{Pm}_{98,100}$. The 3-qp isomers involve one quasi-proton plus two quasi-neutrons near the Fermi surface as shown in Fig. 1. A calculation employing the projected shell model (PSM)⁵⁾ was performed to understand the isomers. The observation of the $(17/2+)$ isomers cannot be explained by the PSM calculation with traditional Nilsson parameters. We demonstrate that to explain the observation, the strength of the spin-orbit interaction needs to be changed according to the neutron number in exotic nuclei, as suggested by Liu *et al.*⁶⁾ Furthermore, we also confirmed that the reported deformed shell gap at $N = 98$ ³⁾ is present both in even-even and odd-mass nuclei.

References

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^{*1} Center for Nuclear Study, University of Tokyo
^{*2} Department of Physics and Astronomy, University of Tennessee, Knoxville
^{*3} Research Center for Nuclear Physics, Osaka University
^{*4} LPSC, Université Grenoble-Alpes, CNRS/IN2P3
^{*5} School of Physics and Astronomy, Shanghai Jiao Tong University
^{*6} School of Physics, Huzhou University
^{*7} School of Physical Science and Technology, Southwest University, Chongqing
^{*8} RIKEN Nishina Center
^{*9} Department of Physics, Tohoku University
^{*10} Department of Physics, Peking University
^{*11} Department of Physics, University of Tokyo
^{*12} Department of Physics, University of Oslo
^{*13} INFN, Sezione di Milano
^{*14} School of Computing Engineering and Mathematics, University of Brighton
^{*15} Department of Physics, Osaka University
^{*16} Dipartimento di Fisica dell'Università degli studi di Padova
^{*17} INFN, Laboratori Nazionali di Legnaro
^{*18} Institut Laue-Langevin, Grenoble
^{*19} Department of Physics, Rikkyo University
^{*20} GSI Helmholtzzentrum für Schwerionenforschung GmbH
^{*21} Department of Physics, University of Tsukuba
^{*22} RISP, Institute for Basic Science, Daejeon
^{*23} Institute for Nuclear Research (ATOMKI)
^{*24} Department of Display Engineering, Hoseo University
^{*25} Department of Physics, University of Surrey
^{*26} Department of Physics, University of York
^{*27} Instituto de Estructura de la Materia, CSIC
^{*28} Departamento de Física Teórica, Universidad Autónoma de Madrid
^{*29} IRCNPC, Beihang University