

## Progress on the analysis of $P_n$ -values relevant for the formation of the $r$ -process rare-earth peak

M. Pallàs,<sup>\*1</sup> A. Tarifeño-Saldivia,<sup>\*2,\*1</sup> G. G. Kiss,<sup>\*3</sup> J. L. Tain,<sup>\*2</sup> A. Tolosa-Delgado,<sup>\*4,\*2</sup> A. Vitéz-Sveiczler,<sup>\*3,\*5</sup> F. Calviño,<sup>\*1</sup> J. Agramunt,<sup>\*2</sup> A. Algora,<sup>\*2</sup> N. T. Brewer,<sup>\*8,\*6</sup> R. Caballero-Folch,<sup>\*9</sup> T. Davinson,<sup>\*10</sup> I. Dillmann,<sup>\*9,\*11</sup> A. Estrade,<sup>\*12</sup> N. Fukuda,<sup>\*7</sup> R. K. Grzywacz,<sup>\*8,\*6</sup> O. Hall,<sup>\*10</sup> N. Mont-Geli,<sup>\*1</sup> A. I. Morales,<sup>\*2</sup> A. Navarro,<sup>\*1</sup> N. Nepal,<sup>\*7</sup> S. Nishimura,<sup>\*7</sup> B. C. Rasco,<sup>\*8,\*6</sup> K. P. Rykaczewski,<sup>\*6</sup> T. N. Szegedi,<sup>\*3</sup> V. Phong,<sup>\*7</sup> R. Yokoyama,<sup>\*13</sup> M. Wolińska-Cichocka,<sup>\*14</sup> and P. J. Woods<sup>\*10</sup> for the BRIKEN Collaboration<sup>\*15</sup>

Rapid neutron capture (the  $r$ -process) produces nearly half of the nuclei heavier than iron in explosive stellar scenarios. Above the mass number  $A = 100$ , there are two main peaks in the  $r$ -process solar-system abundances, are located at  $A \sim 130$  and  $A \sim 195$ . Located between them, the Rare-Earth Peak (REP) is a tiny but definite peak at mass number  $A \sim 160$  that results from the freeze-out during the last stages of neutron exposure. According to theoretical models and sensitivity studies, half-lives ( $T_{1/2}$ ) and  $\beta$ -delayed neutron emission probabilities ( $P_{xn}$ ) of neutron-rich nucleus, in the mass region  $A \sim 160$  for  $55 \leq Z \leq 64$  are critical for the formation of the REP.<sup>1,2)</sup> As a part of the BRIKEN collaboration, the NP1612-RIBF148 experiment measured half-lives and  $\beta$ -delayed from Ba to Gd ( $Z = 56-64$ ). In 2018, an experimental run centered on  $^{165}\text{Pm}$  was conducted using a total of 5 days of beamtime. The results for Pm to Gd ( $Z = 61-64$ ) species are already published.<sup>3)</sup> The data analysis for Ba to Nd ( $Z = 56-60$ ) species using the method proposed in<sup>4)</sup> is presented in this report.

The NP1612-RIBF148 experimental setup consisted of the Advanced Implantation Detector Array (AIDA)<sup>5)</sup> and the BRIKEN neutron counter.<sup>6)</sup> The neutron counter was placed surrounding AIDA to detect the  $\beta$ -delayed neutrons offering a nominal value for the efficiency of 68.6% up to 1 MeV.

In Fig. 1, the preliminary results for the  $P_{1n}$  values are compared with evaluated nuclear data from the ENSDF and some theoretical predictions.<sup>7-9)</sup> Preliminary reports include 14 new  $P_{1n}$  values ( $^{151-152}\text{Ba}$ ,

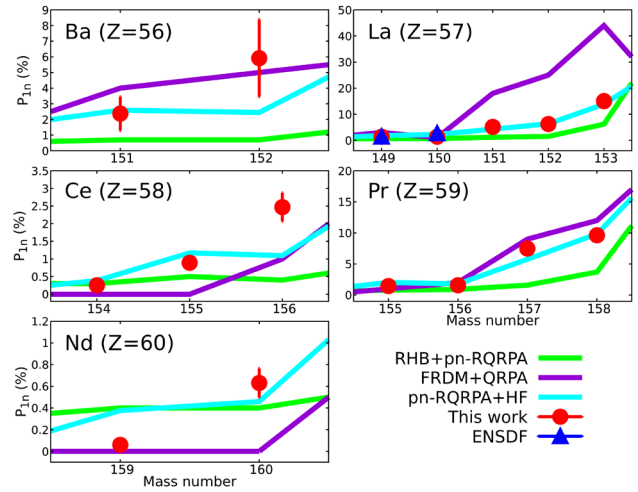


Fig. 1. Preliminary  $P_{1n}$  values derived in this work (red dot) compared with previous measurements (blue triangle) and theoretical calculations.<sup>7-9)</sup>

$^{151-153}\text{La}$ ,  $^{154-156}\text{Ce}$ ,  $^{155-158}\text{Pr}$ , and  $^{159-160}\text{Nd}$ ). Two other  $P_{1n}$  values ( $^{149-150}\text{La}$ ) were also remeasured, obtaining consistent results with previous measurements. Our data supports the overall trend for all  $P_{1n}$  predictions when compared to theoretical models. The  $pn$ -RQRPA + HFM model<sup>9)</sup> best replicates the experimental data and provides good agreement for the majority of the isotopes.

### References

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<sup>\*1</sup> Institut de Tècniques Energètiques (INTE), Universitat Politècnica de Catalunya (UPC)

<sup>\*2</sup> Instituto de Física Corpuscular (IFIC), CSIC-UV

<sup>\*3</sup> Institute for Nuclear Research (ATOMKI)

<sup>\*4</sup> Department of Physics, University of Jyväskylä

<sup>\*5</sup> Doctoral School of Physics, University of Debrecen

<sup>\*6</sup> Physics Division, Oak Ridge National Laboratory

<sup>\*7</sup> RIKEN Nishina Center

<sup>\*8</sup> Department of Physics and Astronomy, University of Tennessee

<sup>\*9</sup> TRIUMF, Vancouver

<sup>\*10</sup> School of Physics and Astronomy, University of Edinburgh

<sup>\*11</sup> Department of Physics and Astronomy, University of Victoria

<sup>\*12</sup> Department of Physics and Science of Advanced Materials Program, Central Michigan University

<sup>\*13</sup> Center for Nuclear Study, University of Tokyo

<sup>\*14</sup> Heavy Ion Laboratory, University of Warsaw

<sup>\*15</sup> www.wiki.ed.ac.uk/display/BRIKEN/Home